SMH

Relevance of Climatological Background Error Statistics for meso-scale Data Assimilation

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The linear regression model for background error statistics

The model for background error statistics in the HARMONIE data assimilation was formulated by Berre (2000)

> $X = \mathbf{B}^{1/2} \xi$ $\mathbf{B}^{-1/2} = V D^{-1} F$

where

- $B^{-1/2}$ is the inverse of square-root of the background error covariance,
- *F* is horizontal 2-dimensional Fourier transform from physical grid-point space to spectral space,
- D^{-1} is a de-correlation operator,
- *V* is a vertical transform utilizing the eigenvectors of vertical covariance matrices.

It is assumed that the spectral components for different wavelengths are statistically uncorrelated what significantly simplifies the formulation of background error covariance in spectral space. The price to pay is the homogeneous background error statistics in physical space. The horizontal isotropy is assumed in addition what allows to simplify formulation of background error covariance even more and to represent horizontal covariances via 1D covariance spectra for control variable.

Two different ensemble sets for generation of background error statistics for HARMONIE AROME 2.5

DS - downscaling of coarse resolution operational ECMWF EDA (T399, 91 vertical level, 12 window 4DVAR) **EDA** – high-resolution HARMONIE AROME 2.5 EDA (6h 3DVAR, conventional observations, perturbed observations for ensemble members except for control) + ECMWF EDA on lateral boundaries

Grid-point structure functions for longitudinal wind speed component in east-west direction







The balance operator D is derived in spectral space C =through step-wise $\eta = MH\zeta + \eta_u$ multivariate statistical $(T, P_s) = NH\zeta + P\eta_u + (T, P_s)_u$ regression technique for each $q = QH\zeta + R\eta_u + S(T, P_s)_u + q_u$ wave number component separately

The regression is performed on an ensemble of short range forecasts differences.



Figure 1: Structure functions for the longitudinal component of the wind in the east-weast direction. Horizontal domain average at model level 47 (\approx 900 hPa) for 5 August 2011 06 UTC. Separate curves for forecast lengths +0h, +1h, +2h, +3h, +4h, +5h, +6h, +9h and +12h. The theoretical slope of the structure function $s^{2/3}$ as well as the structure functions derived from aircraft observations are also included. Downscaling (top) and EDA (bottom).

Figure 2: Spectral densities of +1h, +3h, +6h and +12h unbalanced humidity background error as estimated by the ensemble downscaling (top) and the Ensemble Data Assimilation (EDA, middle) techniques, Model level 35. Spectral densities of +12h unbalanced humidity background error as estimated by the downscaling and EDA techniques (bottom).

DS based structure functions are associated with a strong spinup during first 6-12 hours of integration. The adaptation to a highresolution orography is likely to be one of the main processes behind the spinup. The use of low-resolution ensemble for the purposes of highresolution data assimilation will always be associated with such interpolation processes. One can see that for humidity fields for example certain spin-up processes are on-going even after 12 h of integration with HARMONIE AROME

Climatological Background Error Covariance Model as a diagnostic tool

In the plots to right we show a percentage of the explained **a** surface pressure variance computed from 3 different ensemble sets. The plot **a**) curves correspond to statistics obtained from **BRAND** ensemble and from HARMONIE AROME 2.5 EDA conv **6h3DVAR**; conventional observing network only). The plot **b**) curves correspond to statistics obtained again from **BRAND** ensemble and from HARMONIE AROME 2.5 EDA MetCoOP (3h3DVAR, high resolution observing network including radar data, GNSS, ATOVS). Structure functions are computed from +03h forecast differences for three cases. First of all, we can notice that the gravity wave signature (a strong correlation between mass-field and divergence at scales < 25km) is much stronger for **3h3DVAR** than **6h3DVAR**. This indicates that an initialisation procedure has to be applied even for a highresolution data assimilation. Although the model integration manages to reduce the level of noise, the existence of gravity waves diminishes efficiency of the data assimilation procedure



Alternative ensemble generation approach



The Scheme: generation of perturbations with the structure of B-matrix covariance.



Secondly we can notice a pick in the "Pb" curve at scales 200-300 km when statistics are computed from EDA ensembles, a high narrow pick in case of 6h3DVAR EDA conv and a smoother lower pick in case of 3h3DVAR EDA MetCoOp. Such a pick does not appear in statistics computed from the **BRAND** ensemble. We claim that such a "pick" is a result of the observation perturbation methodology for ensemble generation. Properties of a short range ensemble reflects the way in which an ensemble was generated. In 6h3DVAR EDA conv the random unstructured perturbations were induced at the scale corresponding to the scales of conventional observing system. The background error covariance model is used to transform this unstructured noise into structured perturbations. The model integration will transfer perturbations injected at particular scales to the entire model space. However as we see from plot c) to the left, most of the injected energy will be transferred to larger scales.

BRAND - ensemble perturbations are generated in control vector space and projected to the physical space applying square-root of B-matrix covariance. The perturbations are added to the short range control forecast and relaxed on the LBC's towards ECMWF EDA.

BRAND perturbations perturb entire model space in a structured way; they impose structures as described by the background error covariance model, the same model that is used to form the analysis increments from observations.

EDA perturbations inject random unstructured spread on the scales of observing system. EDA relies on DA technique and the model integration to spread the injected energy in the structured way to entire model space.

Filtered analysis increment

A typical HARMONIE AROME 2.5 3h3DVAR analysis temperature increment is shown in plot to the right. One can clearly see large scale structures with superimposed small scale noise. This agrees well with the diagnostics above based of the statistical forecast error structures. This small scale noise is due to gravity waves present in the ensemble from which the structure functions were derived. We claim that the gravity waves do not reflect the genuine nature of meso-scale flow but rather appear as artefacts of ensemble generation methodology lacking the initialisation step. Below we show the comparison of the verification scores between a 1 month reference run (standard **3h3DVAR**; red) and the "filtered analysis increment" experiment (**3h3DVAR 100**; green) where the only 100 longest waves (>25km) were



On the basic assumptions behind climatological background error covariance model

The model for the HARMONIE background error statistics is based on the three assumptions of stationarity, homogeneity and isotropy with respect to horizontal covariances. A very large size ensemble is needed to obtain a stable model for background error statistics when it is derived statistically from the ensemble of model differences. Horizontal averaging (through assumption of the horizontal homogeneity) is a very efficient remedy to obtain stable and smooth background error statistics. The assumption of horizontal isotropy adds even more smoothing. Although such a model may be attractive from a computational efficiency point of view, one may ask how representative stationary, isotropic and homogeneous horizontal correlations are for the forecasts with a strong caseto-case variability. Our experience shows that variational data assimilation at convective scale profits a lot even if a relatively simple ensemble is used to describe consistently error-of-the-day (for example via a Hybrid formulation). This is probably due to the fact that convective motion happens in the low part of atmosphere which is strongly affected by orographic conditions and even a relatively small ensemble size is enough to capture inhomogeneity and anisotropy associated with orography.









In the plots to the right the balance operator **D** derived from **BRAND** and **3h3DVAR EDA MetCoOp** are presented. MetCoOp observing network is heavily dominated by mass observations. The geostrophic balance imposed by the background error covariance model is clearly seen in the structure functions derived from 3h3DVAR EDA MetCoOp data set.

